# Lecture 5 Notes: Stellar Structure and Flux Measurements

Summary Notes

Aug-Sep 2024

### 1 Learn programming

- Python programming recommended for astronomy
- Reference: https://web.iucaa.in/ ace/programs/

### 2 Fluxes and Magnitudes

#### 2.1 Apparent Magnitude

- Basic equation:  $m_1 m_2 = -2.5 \log_{10}(f_1/f_2)$
- Key characteristics:
  - Logarithmic scale
  - Base  $100^{1/5}$
  - Inverted (brighter objects have lower magnitude)
  - Measurements are relative
  - Small flux changes: fractional change in flux  $\sim$  magnitude change

#### 2.2 Magnitude Systems

- 1. Vega system (historical standard)
- 2. STMAG system (HST standard)
  - $m_{\lambda}(\lambda) = -2.5 \log_{10} F_{\lambda}(\lambda) 21.1$
  - $F_{\lambda}$  in erg s<sup>-1</sup> cm<sup>-2</sup> Å<sup>-1</sup>
- 3. AB magnitude system
  - $m_{\nu}(\lambda) = -2.5 \log_{10} F_{\nu}(\lambda) 48.6$
  - $F_{\nu}$  in erg s<sup>-1</sup> cm<sup>-2</sup> Hz<sup>-1</sup>

#### 2.3 Jansky Units

- 1 Jy =  $10^{-23}$  erg/sec/Hz/cm<sup>2</sup>
- Conversion:  $f = 3631 \text{ Jy} \times 10^{-0.4m}$
- Common in radio and far-infrared astronomy

### **3** Distance Measurements

#### 3.1 Distance Modulus

- $\mu = m M = 5 \log_{10}(D_L) 5$
- $D_L$  is luminosity distance in parsecs
- Interpretation:
  - $-\mu > 0$ : object farther than 10 parsecs
  - $\mu < 0$ : object closer than 10 parsecs

### 4 Stellar Structure and Evolution

#### 4.1 Definition of a Star

Two key conditions:

- 1. Bound by self-gravity
- 2. Powered by internal energy source

#### 4.2 Hydrostatic Equilibrium

- Balance between gravity and pressure
- Key assumptions:
  - No magnetic field
  - Non-rotating star
  - General relativity not significant

#### 4.3 Eddington Luminosity

- Maximum luminosity while maintaining hydrostatic equilibrium
- Classical formula:  $L_{\rm Edd} = \frac{4\pi G M m_p c}{\sigma_T}$
- Modified by opacity:  $L_{\rm Edd,mod} = \frac{4\pi GMc}{\kappa}$

- Applications:
  - Limits on stellar masses
  - Accretion rates onto black holes
  - AGN feedback mechanisms

## 5 Energy Transport in Stars

Three main mechanisms:

- 1. Conduction
- 2. Radiation
- 3. Convection (most complex mechanism)

### 6 Equations of State

- Ideal gas: P = nkT/V
- Non-relativistic degenerate electron gas:  $P = K(n/V)^{5/3}$
- Relativistic degenerate electron gas:  $P = K'(n/V)^{4/3}$
- Polytropic equations:  $P = K \rho^{\gamma} = K \rho^{1+1/n}$

### 7 Nuclear Fusion

#### 7.1 Proton Chain

Net reaction:  $4^1H \rightarrow {}^4He + 2\nu_e + 26.7 \text{ MeV}$ 

- Sun produces  $4 \times 10^{26}$  joules per second
- Approximately  $9.36 \times 10^{37}$  reactions per second
- Efficiency around 0.7%

### 8 Solar Composition

Mass fractions:

- Hydrogen (X): 0.70
- Helium (Y): 0.28
- Metals (Z): 0.02