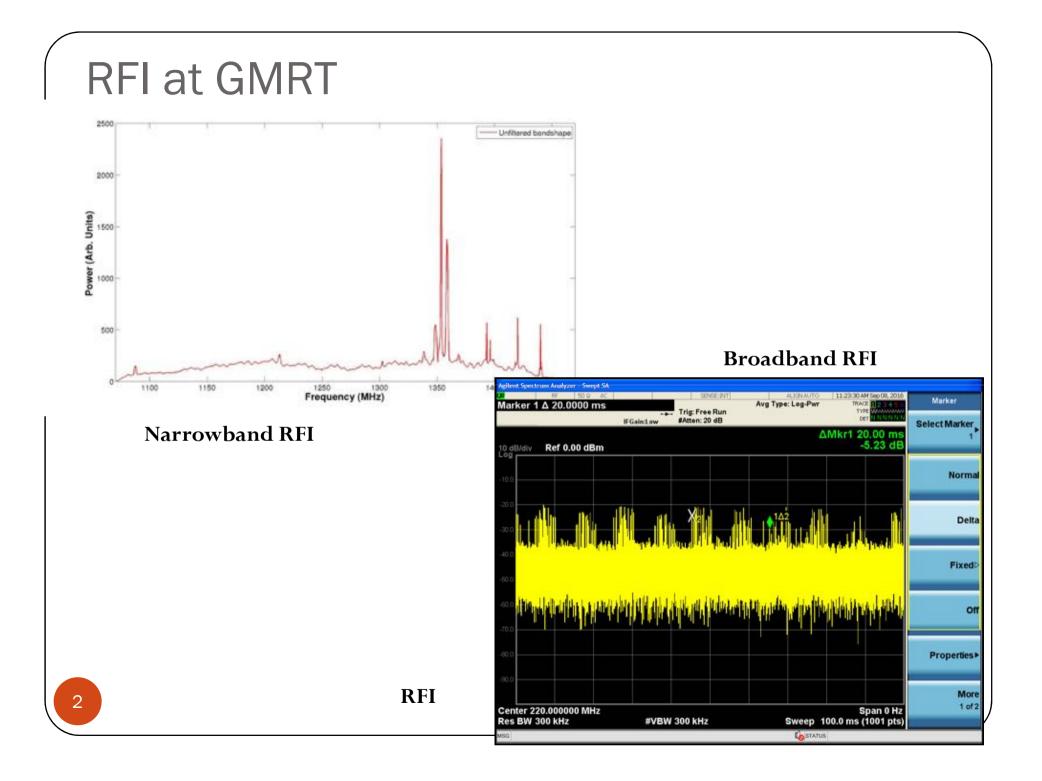
## Results from Real-time RFI Excision System for uGMRT

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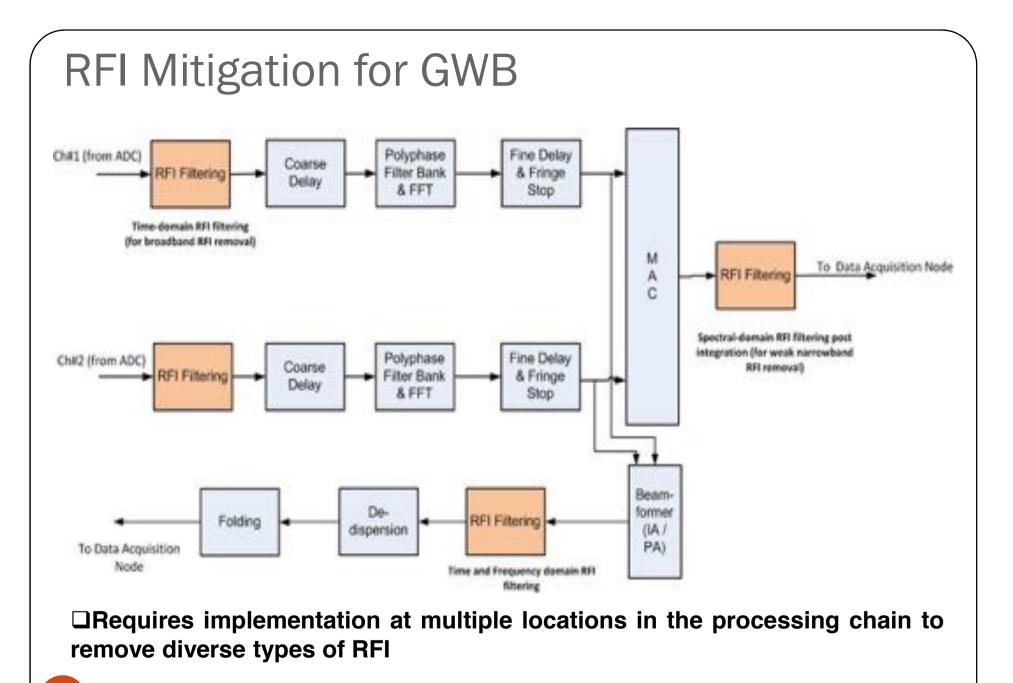
### Why Real-time ?

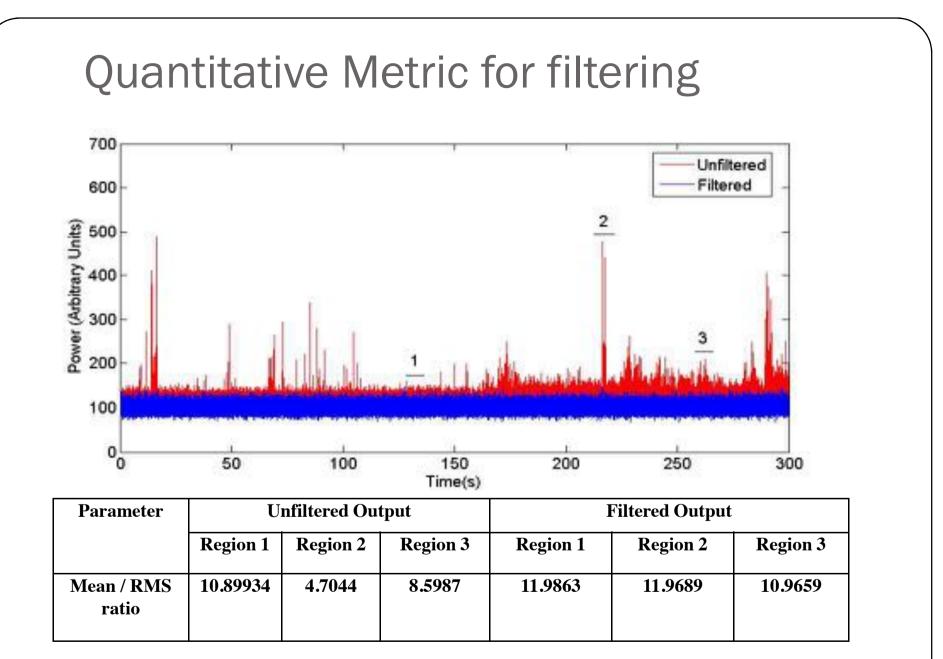
- Temporally impulsive RFI: Energy spreads post-FFT hence excision is needed before FFT.
  - Power-line RFI: Low duty cycle but high spectral occupancy
  - RFI is correlated in closely spaced antennas
- Spectrally impulsive RFI: RFI excision useful for low time occupancy
- Best possible time resolution: reduction in loss of astronomical data due to flagging
- Leads to improvement in receiver sensitivity

#### A stitch in (real) time saves nine !

## **RFI** Excision

- RFI in astronomical data outliers make Gaussian distribution heavy-tailed
- Excision assumes that RFI is much stronger than the astronomical signal
- Robust threshold using Median Absolute Deviation for RFI detection
- Excision by replacing the RFI affected samples by constant value or noise or threshold
  - Implemented in temporal and spectral domains





#### Theoretical: V(B\*T) = V((200e6/2048)\*1.31ms) = 11.31

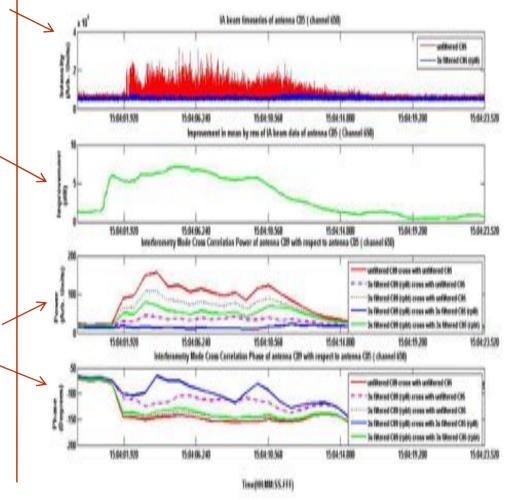
• High time resolution:

Single spectral channel (650 MHz) plot over time (IA mode) from the GWB at 1.3 ms time resolution for filtered and unfiltered outputs

- Improvement (dB)
  - $I = 10log(MR_F/MR_U)$

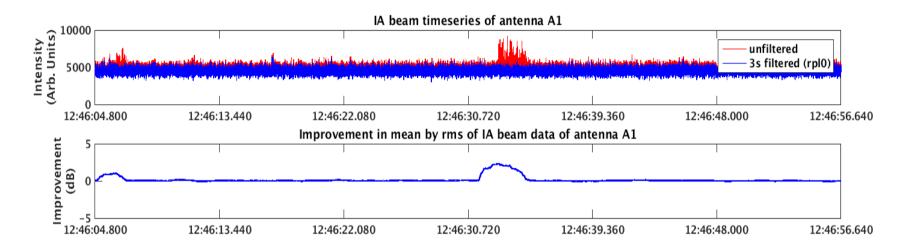
where MR\_F and MR\_U are the mean/rms ratio for filtered and unfiltered signal respectively. Running mean/rms calculated over 1024 samples of IA beam output

Coarse time resolution (671 ms): Cross-correlation magnitude (unnormalized) and phase – options – filtered vs filtered, and unfiltered vs unfiltered for short baseline



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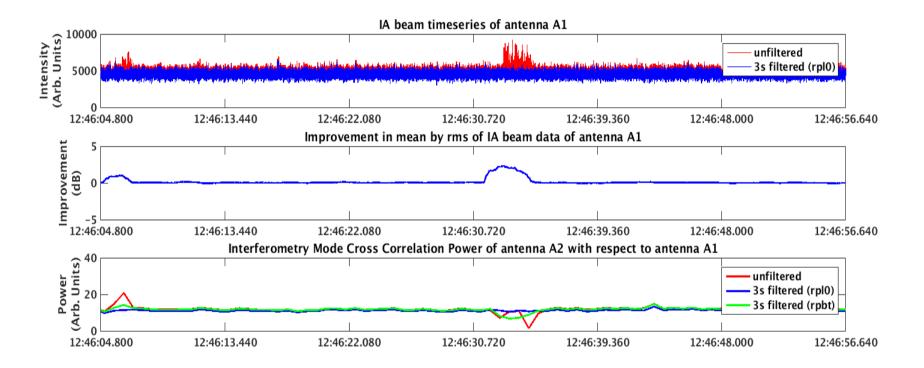
## Test Results (Antenna signals)



8

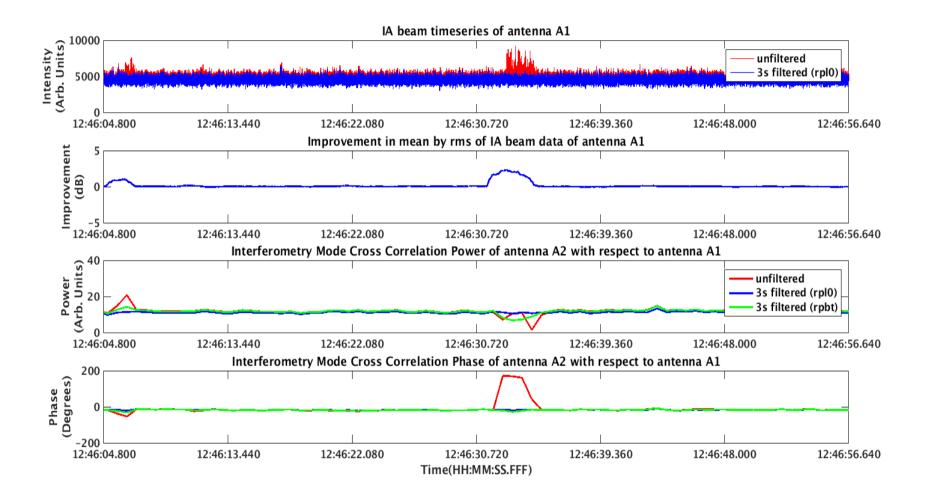
Beam output of a spectral channel showing filtering at 3σ threshold – replacement with zero and threshold

## Test Results (Antenna signals)

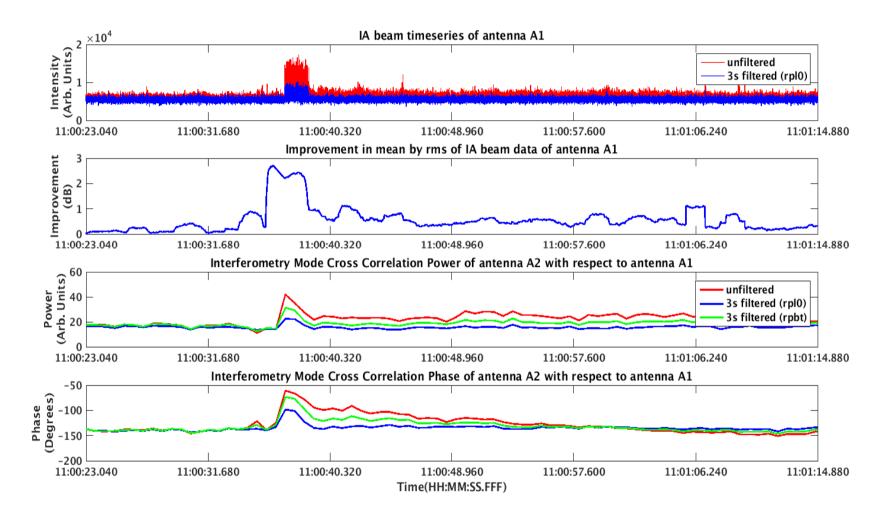


Beam and correlator output of a spectral channel showing filtering at 3σ threshold – replacement with zero and threshold

## Test Results (Antenna signals)

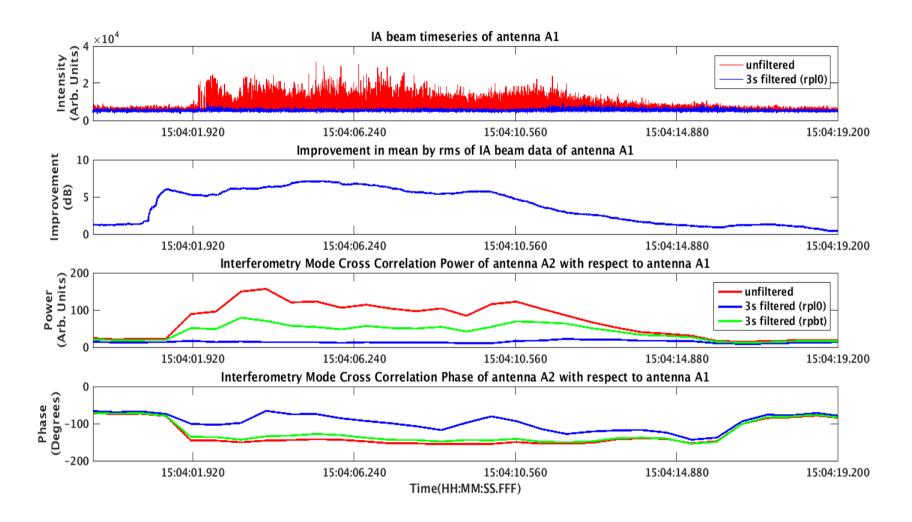


Beam and correlator data of a spectral channel showing filtering at 3σ threshold – replacement with zero and threshold

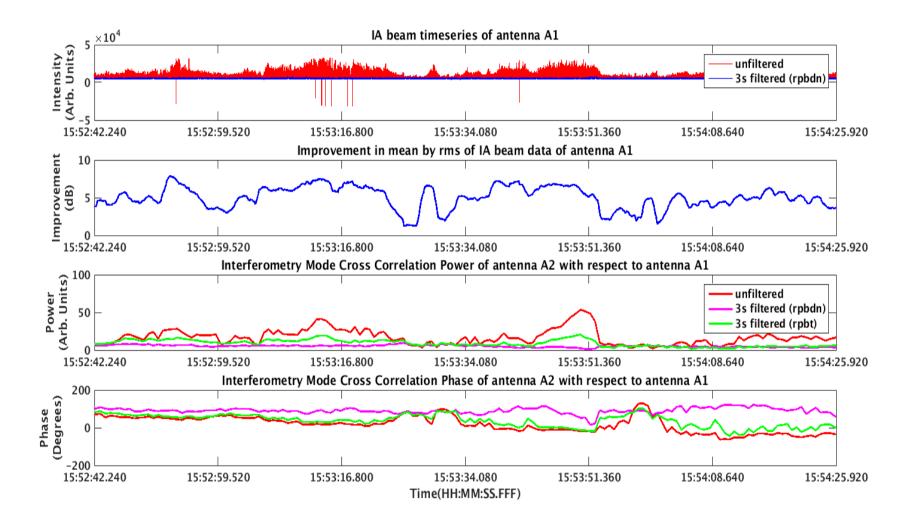


Beam and correlator data of a spectral channel showing filtering at 3σ threshold – replacement with zero and threshold

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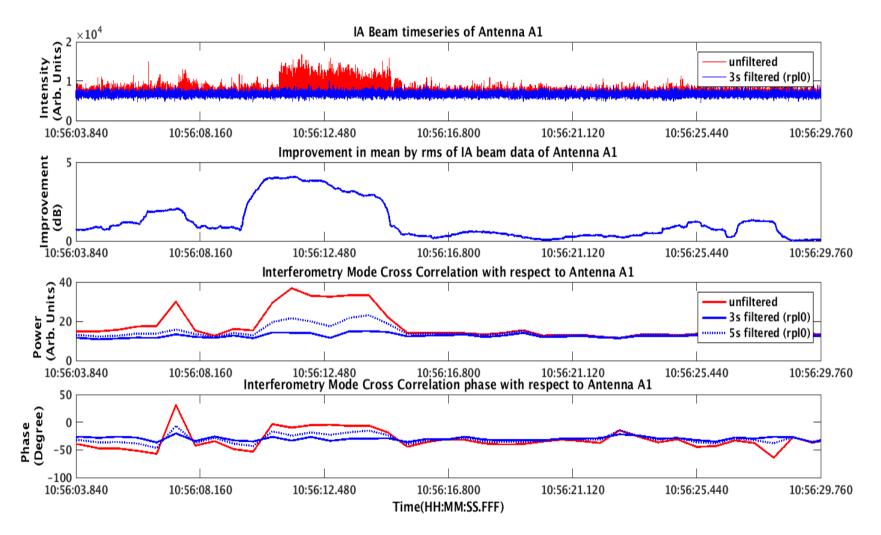


Beam and correlator data of a spectral channel showing filtering at 3σ threshold – replacement with zero and threshold

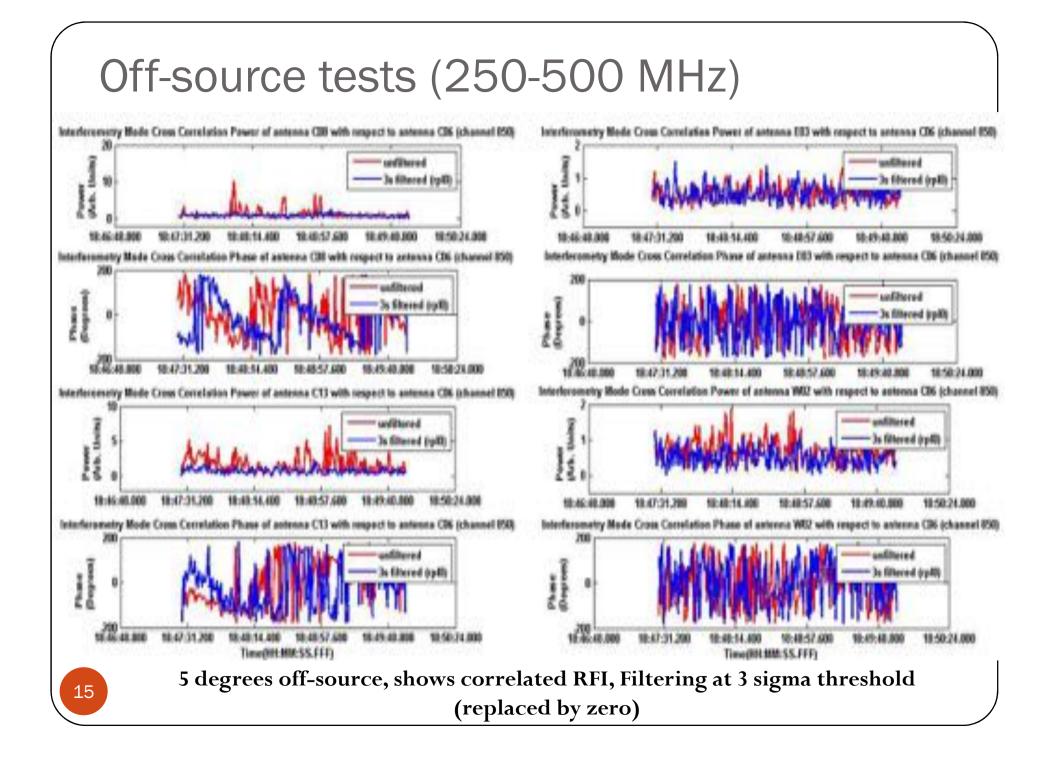


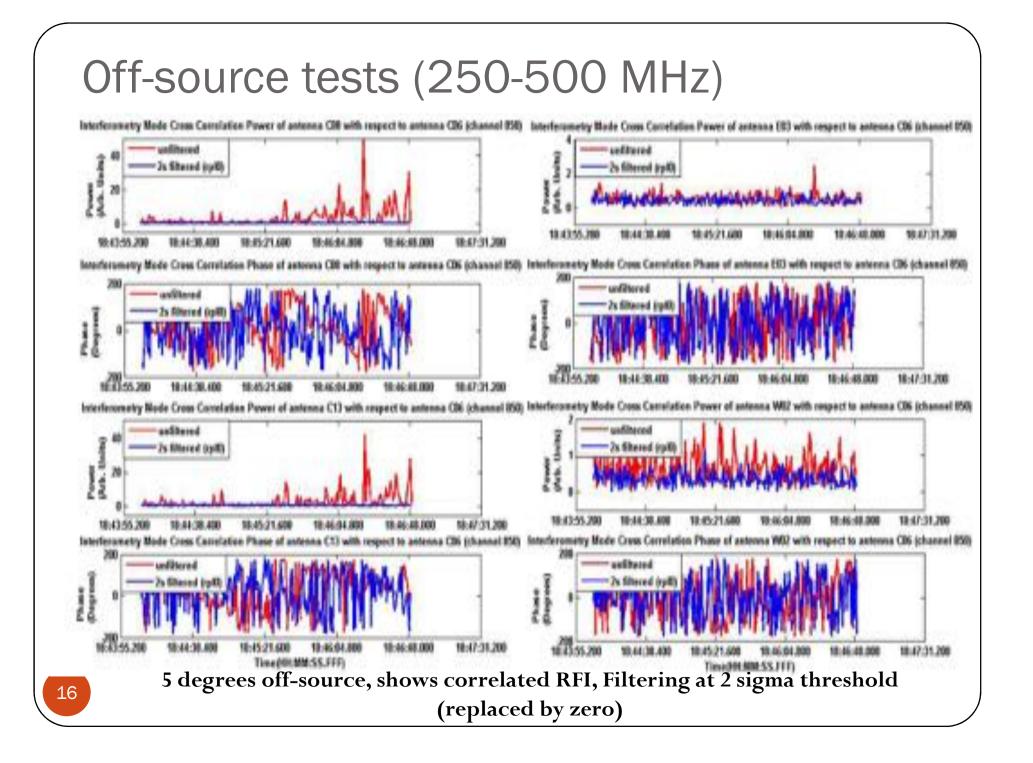
Beam and correlator data of a spectral channel showing filtering at 3σ threshold – replacement with zero and noise

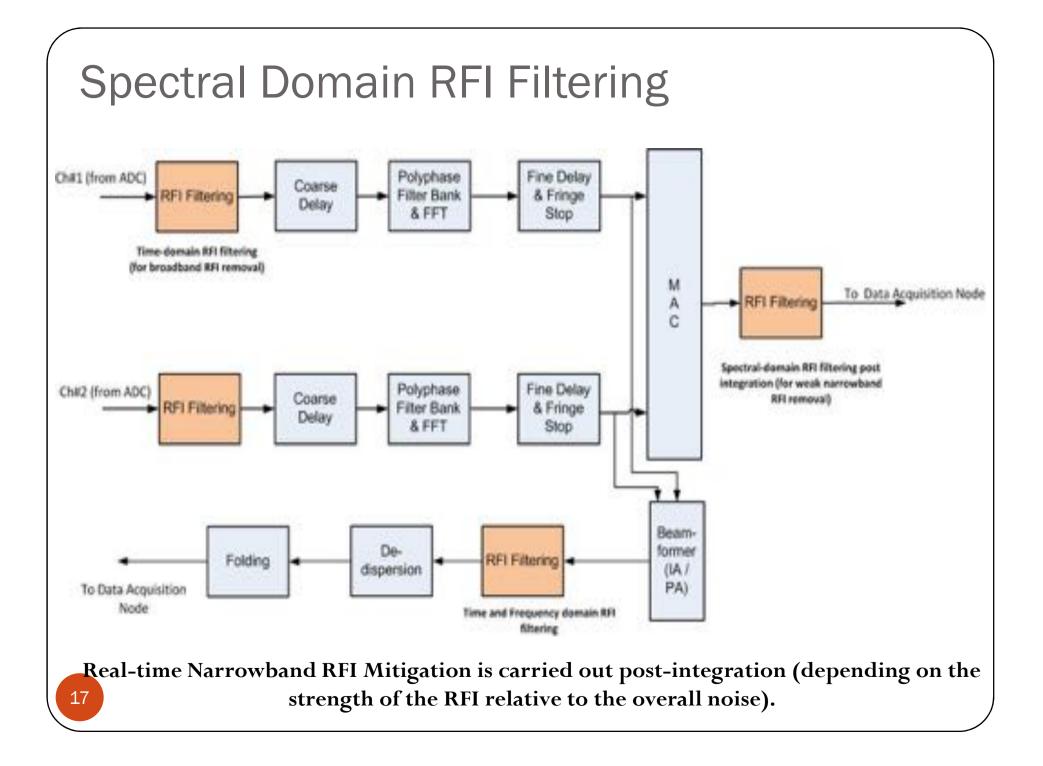
## Test using different thresholds



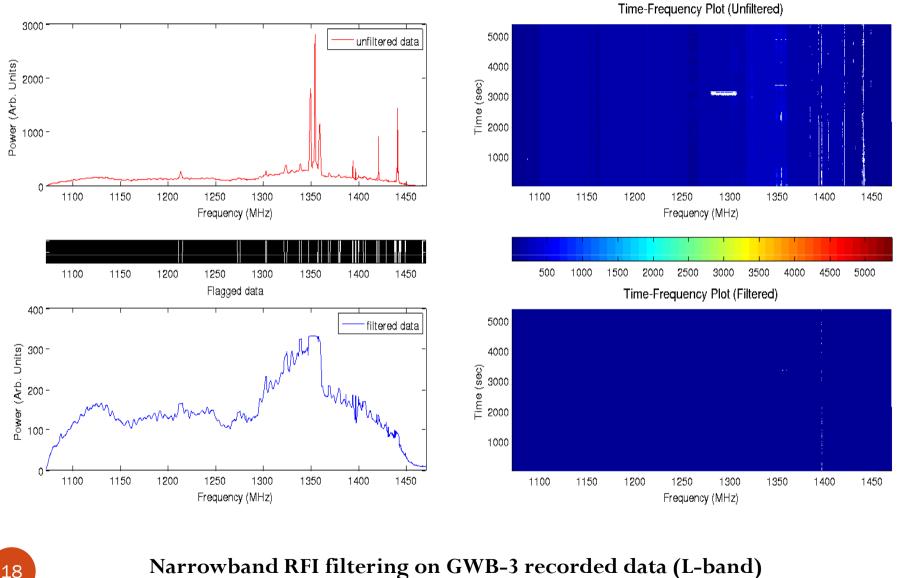
Beam and correlator data of a spectral channel showing filtering at  $3\sigma$ and  $5\sigma$  threshold – replacement with zero





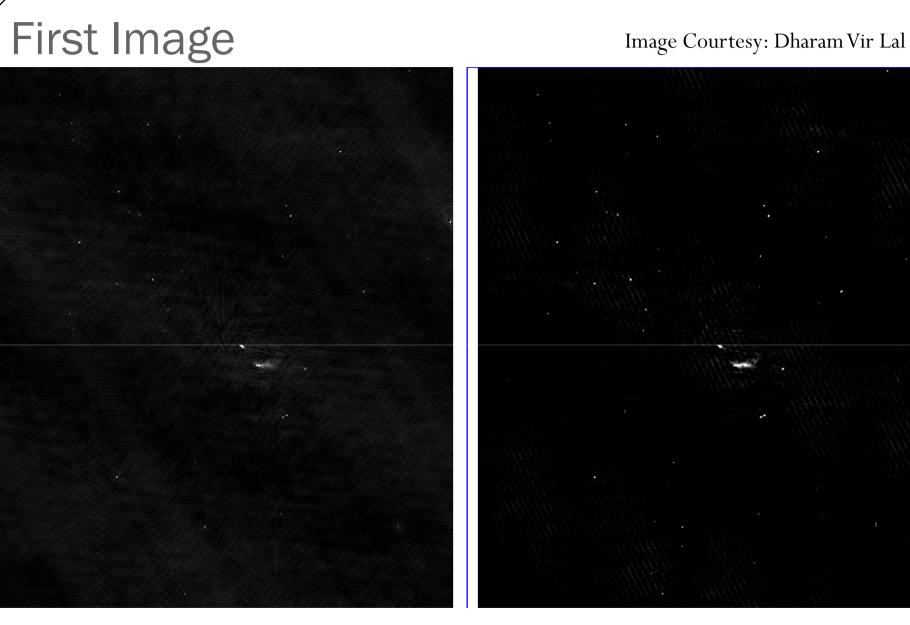


## Narrowband filtering on GWB data



5000 s data single-antenna plot

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250-500 MHz, 16 antennas, Ch-1 (left, without filter) & Ch-2 (right, with broadband RFI filter), factor of two improvement post-filtering

### Real-time RFI Filtering: Features

- Possible to set fractional filtering threshold (in steps of 0.1)
- Various filtering options available (constant values, threshold, digital noise)
- Filtering possible in Ch-1 or Ch-2 or both
- 'Bypass'option is also available
- Filtering for 200 or 400 MHz mode as well as narrowband modes of GWB-3
- Number of samples flagged per antenna for a given scan
  - % RFI can be calculated as this feature provides total number of samples and the number of flagged samples

## Summary

- Broadband RFI mitigation using MAD-based filtering in realtime is available for the uGMRT user community.
- Various tests carried so far show an improvement of 10-12 dB in the post-filtering signal-to-noise ratio.
- Long-term RFI filtering tests to understand the effect of filtering on power spectrum and cross-correlation are being carried out
- Narrowband RFI mitigation has been demonstrated on recorded visibilities.

#### Acknowledgements

Shruti Bhatporia Kishor Naik Sanjay Kudale

**GMRT Backend Team** 

**GMRT** Control Room

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# **Thank You!**