Introduction to Astronomy and Astrophysics I Lecture 2

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The Solar System



Credit: Universe Today

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- The Solar System: our cosmic neighborhood
- Composed of the Sun, planets, moons, asteroids, comets, and dust
- Formed approximately 4.6 billion years ago How do we know this?
- Gravitationally bound system
- Extends from the Sun to the outer edges of the Oort Cloud
- Divided into inner and outer regions

- Central star of our Solar System
- Accounts for 99.86% of the system's mass
- Main-sequence star (G-type)
- Fusion reactor: converts hydrogen to helium, 4 million tons of mass is lost every second

- Diameter: 1.4 million km (109 times Earth's diameter)
- Surface temperature: 5,800 K
- Core temperature: 15 million K
- Drives the Solar System's dynamics and energy

Anatomy of the Sun

ANATOMY OF THE SUN

Sunspots

Darker, cooler areas on the photosphere with concentrations of magnetic field

Prominence Large structure, often many thousand of kilometres in extent

Granulation

Small, short-lived grainy features tha cover the Sun, caused by thermal currents rising from below

Chromosphere

Layer above the photosphere, where the density of plasma drops dramatically

Photosphere The visible 'surface' of the Su

Transition region Thin, irregular layer that separates the relatively cool chromosphere from the much hotter corona

Flare Sudden release of energy in th form of radiation

#SolarOrbiter #WeAreAllSolarOrbiters

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Convective zone Rapid heating of plasma creates currents of heated and cooled gas

Radiative zone Energy created in the core diffuses slowly through the plasma

Core

Where the Sun generates its energy via thermonuclear reactions

- Corona

The Sun's outer atmosphere which extends millions of kilometres into outer space

Coronal mass ejection

Vast eruption of billions of tonnes of plasma and accompanying magnetic fields from the Sun's corona

Solar wind

A continuous stream of charged particles released from the corona



Resolving the "coronal heating" problem

- Wave heating: Magnetohydrodynamic (MHD) waves, particularly Alfvén waves, propagate from the lower layers of the Sun's atmosphere These waves carry energy and dissipate in the corona, heating it.
- Magnetic reconnection: The Sun's magnetic field lines can become tangled and then snap back into place. This process, called magnetic reconnection, releases significant energy. Nanoflares, small but frequent reconnection events, may contribute to consistent heating
- Alfvén wave turbulence: As Alfvén waves propagate, they can become turbulent. This turbulence can cascade to smaller scales where the energy is more easily dissipated as heat.

The exact combination and relative importance of these mechanisms is still an active area of research using Parker Solar Probe, Aditya L1, Solar Orbiter.

- Visible Emission Line Coronagraph (VELC): For studying the solar corona and origin of CMEs
- Solar Ultraviolet Imaging Telescope (SUIT): For imaging the photosphere and chromosphere
- Aditya Solar Wind Particle Experiment (ASPEX): For studying solar wind properties
- Plasma Analyser Package for Aditya (PAPA): For solar wind ion analysis
- Solar Low Energy X-ray Spectrometer (SoLEXS): For studying solar flares
- High Energy L1 Orbiting X-ray Spectrometer (HEL1OS): For studying solar flares in higher energy X-rays
- Magnetometer: For measuring interplanetary magnetic field

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- The corona is much less dense than the chromosphere. Lower density means that the same amount of energy input heats the corona more effectively
- The magnetic field dominates the plasma dynamics in the corona, making magnetic heating mechanisms more efficient

- Also known as terrestrial or rocky planets
- Include Mercury, Venus, Earth, and Mars
- Characterized by:
 - Solid, rocky surfaces
 - Relatively small sizes
 - Few or no moons
 - No ring systems

Inner Planets Comparison



Credit: Wikipedia

- Smallest planet in the Solar System
- Closest to the Sun
- No atmosphere, extreme temperature variations
- Heavily cratered surface
- Unexpected magnetic field

Mercury Surface





- Similar in size to Earth
- Thick atmosphere of CO₂
- Extreme greenhouse effect
- Surface temperature: 462 °C
- Retrograde rotation 243 days to turn on its axis and 224.7 days to orbit the Sun

Venus Atmosphere



is the fraction of light reflected by a planet. It is 76% for Venus, 30% for Earth and 7% for the Moon.

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- Third planet from the Sun
- Only known planet with life
- 71% of surface covered by water
- Nitrogen-oxygen atmosphere
- Active geology and weather systems

Earth from Space



Credit: Encyl. Britannica

- Largest moon relative to its planet's size in the Solar System
- Synchronous rotation with Earth
- Influences Earth's tides and stabilizes its axial tilt
- Formation likely due to giant impact early in Earth's history

- Fourth planet from the Sun
- Known as the "Red Planet"
- Thin atmosphere, mostly CO₂
- Evidence of past water activity
- Two small, irregular moons: Phobos and Deimos

Mars Surface Features



- Region between Mars and Jupiter
- Contains millions of asteroids
- Ranges from dust-sized to dwarf planet Ceres (940 km diameter)
- Likely remnants from the Solar System's formation
- Source of most meteorites on Earth

Asteroid Belt



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- Also known as gas giants or Jovian planets
- Include Jupiter, Saturn, Uranus, and Neptune
- Characterized by:
 - Large size and mass
 - Primarily composed of hydrogen and helium
 - Numerous moons and ring systems
 - No solid surface (except possibly rocky/icy cores)

Outer Planets Comparison



- Largest planet in the Solar System
- Mass 2.5 times that of all other planets combined
- Great Red Spot: a giant storm raging for centuries
- Powerful magnetic field
- At least 79 known moons

Jupiter and its Moons





- Jupiter: largest planet in our Solar System, Mass: 317.8 Earth masses
- Brown dwarfs: objects between planets and stars, Minimum mass: \sim 75 Jupiter masses (\sim 0.013 solar masses)
- Jupiter's mass: much below the brown dwarf threshold
- Key difference: nuclear fusion
 - Brown dwarfs can fuse deuterium
 - Jupiter cannot initiate any nuclear fusion

Jupiter radiates more energy than it receives from the Sun. How is this possible?

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Jupiter radiates more energy than it receives from the Sun. How is this possible?

Jupiter emits almost twice as much energy as it receives from the sun due to its internal heat source, which is powered by its slow gravitational collapse. Astronomers estimate that Jupiter is shrinking by almost 2 centimeters per year. Jupiter's internal structure is believed to consist of an outer mantle of fluid metallic hydrogen and a denser inner core. As helium rain falls from the top of the metallic hydrogen layer towards the center of the planet, it generates kinetic energy.

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- Second-largest planet
- Known for its extensive ring system
- Lowest density of any planet (less than water)
- At least 82 known moons, including Titan

Saturn's Rings and Aurorae





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- Third-largest planet in the Solar System
- Ice giant: primarily composed of ices (water, ammonia, methane)
- Tilted axis of rotation (97.77°): unique seasonal changes
- Faint ring system
- 27 known moons, named after Shakespearean characters

Uranus and Its Moons



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- Outermost known planet in the Solar System
- Another ice giant, similar composition to Uranus
- Discovered through mathematical predictions
- Great Dark Spot: large storm system
- Strongest winds in the Solar System (up to 2,100 km/h)
- 14 known moons, largest being Triton

Neptune



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- New category created in 2006
- Orbit the Sun and have sufficient mass for gravity to make them nearly spherical
- Have not cleared their orbital neighborhood
- Five recognized: Pluto, Eris, Haumea, Makemake, and Ceres

- Formerly classified as the ninth planet
- Highly eccentric and inclined orbit
- Smaller than Earth's moon
- Has five known moons, largest being Charon
- Visited by New Horizons spacecraft in 2015

Pluto and Charon

