Astronomical Techniques II Lecture 7 - Correlators and Calibration Framework

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Weiner-Khinchin Theorem

•
$$r(\tau) = V_1(t) \bigstar V_2(t) = \lim_{T \to \infty} \frac{1}{2T} \int_{-T}^{T} V_1(t) V_2^{\star}(t-\tau) dt$$

•
$$V_1(t) \bigstar V_2(t) = \frac{1}{2T} \int_{-T}^{T} V_1(t) V_{2-}^{\star}(\tau - t) dt$$

where
$$V_{2-}(t)=V_2(-t)$$

- Now $V_1(t) \rightleftharpoons \hat{V_1}(\nu); V_2(t) \rightleftharpoons \hat{V_2}(\nu); V_{2-}^{\star}(t) \rightleftharpoons \hat{V_2}^{\star}(\nu);$
- From Convolution theorem

$$V_1(t) \bigstar V_2(t) \rightleftharpoons \hat{V}_1(\nu) \ \hat{V}_2^{\star}(\nu)$$

• When $V_2(t) = V_1(t)$, it becomes the Weiner-Khinchin relation

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Weiner-Khinchin Theorem

 Power (density) spectrum of a signal is the FT of its auto-correlation.

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•
$$|V(\nu)|^2 = \int_{-\infty}^{\infty} r(\tau) e^{-i2\pi \nu \tau} d\tau$$

Correlators

- Devices to measure the *mutual coherence function*
- Measuring the cross-correlation function of voltage signals from each of the antennas
- Digital correlators require sampling and quantization

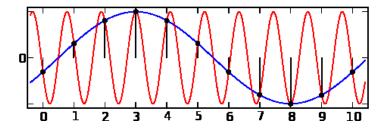
Sampling

 \blacksquare Band limited signal - information limited to a finite bandwidth $\Delta \nu$

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- \blacksquare Baseband signal Mixed down RF signal such that it lies between 0 and $\Delta\nu$
- Minimum sampling frequency = $2\Delta\nu$ (Nyquist criterion)
- Undersampling and Oversampling
- Aliasing

Aliasing

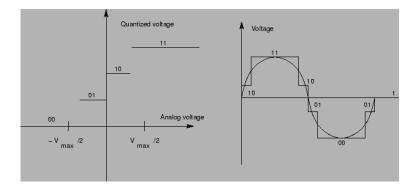


• $\nu_{blue} = 0.1 Hz$; $\nu_{red} = 0.9 Hz$; $\nu_{sampling} = 1 Hz$

- Samples indistinguishable from a signal at frequency $\nu N \times \nu_{sampling}$, where N is an integer
- $N \neq 0$ images or aliases of ν
- Nyquist sampling (sampling at $2\Delta\nu$) prevents aliasing

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Quantization



1
$$V_{Error} = V_{True} - V_{Quantized}$$

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Quantization

- Quantization distorts both the amplitude and the spectrum of the input signal
- **2** Spectrum of quantized signal extends beyond the original $\Delta \nu$ of V_{True} , implies aliasing
- **3** Largest value which can be expressed (within the error of $\pm q/2$) depends on the no. of bits (M) is $q(2^M 1)$
- 4 $V_{Max} = 4.42\sigma$, probability of exceeding V_{Max} is 10^{-5} .

Correlator...

- Dynamic Range minimum change in signal which can be expressed is q
- 2 Discreet Fourier Transform (vs. Continuous Fourier Transform)
 - 1 Windowing, Sampling and Filtering
- **3** Digital delays
 - 1 Discreet delays in units of sampling time
 - **2** A delay of τ correponds to a $\phi = 2\pi\nu\tau$. So delays smaller than the sampling time are corrected by applying phase gradients to the sampled data.

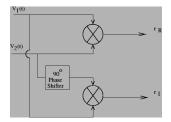
Correlators...

1
$$R(m) = \frac{1}{N} \sum_{n=0}^{N-1} v_1(n) v_2(n+m) \quad 0 \le m \le M$$

- **2** Correlation measured by a digital correlator differs from that measured by an ideal device with infinite precision $(R_c(m))$.
- 3 Deviation depends on the value of correlation and the number of correlator bits - Van Vleck Correction
- Monotonic and approx linear over a large range of correlation values for correlators with a large number of bits (3-4).

Correlators...

1
$$r_R(\tau_g) = Re[v_1(\nu, t) v_2^*(\nu, t)] = |\mathcal{V}| \cos(2\pi\nu\tau_g - \phi_{\mathcal{V}})$$

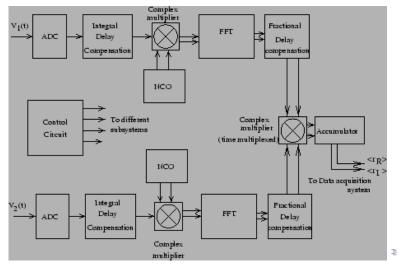


$$\mathbf{I} \ r_{l}(\tau_{g}) = |\mathcal{V}| \ \cos(2\pi\nu\tau_{g} - \phi_{\mathcal{V}} + \pi/2)$$

$$\mathbf{2} \ |\mathcal{V}| = sqrtr_{R}^{2} + r_{l}^{2}; \ \phi_{\mathcal{V}} = tan^{-1}\frac{r_{l}}{r_{R}}$$

Spectral Correlators

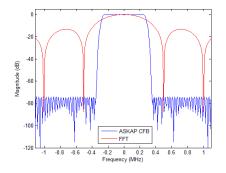
1 FX correlators



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Present/New Generation Correlators

- **1** FFT \rightarrow Polyphase filters
- 2 Real time sample level statistics and data flagging
- 3 Multiple modes time resolution vs spectral resolution



Calibration Framework

$$I V(u,v) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} A_{\nu}(l,m) B_{\nu}(l,m) e^{-2\pi i (ul+vm)} dl dm$$

2
$$V_{i,j}(t) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} A_{\nu}(l,m) B_{\nu}(l,m) e^{-2\pi i (u_{i,j}(t)l + v_{i,j}(t)m)} dl dm$$

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$$\phi_{g} = 2\pi\nu\tau_{g} = 2\pi w = \frac{2\pi}{\lambda} (L_{x} \cos H \cos \delta - L_{y} \sin H \cos \delta + L_{z} \sin \delta)$$

4 Need to know - L_x, L_y, L_z , time, α, δ .

Calibration Framework

G_{i,j}(t) - baseline based complex gain
 ϵ_{i,j}(t) - baseline based complex offset
 eta_{i,j}(t) - gaussian random complex noise

Editing and Flagging

Getting rid of data *known* to be bad
 Getting rid of data *ascertained* to be bad

Calibration Methods

- 1 Direct Calibration
- **2** Sky based calibration (calibrator sources)
- 3 Self-calibration

Antenna based calibration

1
$$\mathcal{G}_{i,j}(t) = g_i(t) \ g_j^*(t) = a_i(t)a_j(t)e^{i(\phi_i(t)-\phi_j(t))}$$

- **2** No. of constraints $\sim N(N-1)/2$
- **3** No. of independent DoF $\sim 2N$
- 4 Vastly over determined problem

- **1** Determining pointing offsets $(\Delta Az, \Delta El)$
- 2 For each antenna and feed
- **3** Sources of error (~ 10 " for GMRT)
 - 1 Tracking errors Servo system feed back loop
 - 2 Distorting of the dish due to gravity Pointing model
 - 3 Wind buffeting